Search for Supersymmetry with Gauge-Mediated Breaking in Diphoton Events with

Missing Transverse Energy at CDF II

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We present the results of a search for supersymmetry with gauge-mediated breaking with $\tilde{\chi}_1^0 \to \gamma \tilde{G}$ in the $\gamma\gamma+$ missing transverse energy final state. In $2.6\pm0.2~{\rm fb}^{-1}$ of $p\bar{p}$ collisions at $\sqrt{s}=1.96~{\rm TeV}$ recorded by the CDF II detector we observe no candidate events, consistent with a standard model background expectation of 1.4 ± 0.4 events. We set 95% C.L. cross section limits and place the world-best limit on the $\tilde{\chi}_1^0$ mass of 149 ${\rm GeV/c^2}$ at $\tau_{\tilde{\chi}_1^0}=0$ ns as well as make exclusions in the $\tilde{\chi}_1^0$ mass-lifetime plane for $\tau_{\tilde{\chi}_1^0}\lesssim 2$ ns.

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The standard model (SM) of elementary particles has been enormously successful, but is incomplete. Theoretical motivations [1] and the observation of the ' $ee\gamma\gamma$ +missing transverse energy (E_T)' [2, 3] candidate event by the CDF experiment during Run I at the Fermilab Tevatron provide compelling rationale to search for the production and decay of new heavy particles that produce events with final state photons and E_T in collider experiments. Of particular theoretical interest are supersymmetry (SUSY) models with gauge-mediated SUSY-breaking (GMSB) [1]. These models solve the "naturalness problem" [4] and provide a low mass dark matter candidate that is both consistent with inflation and astronomical observations [5]. Since many versions of these models have a similar phenomenology, we 11 consider the scenario in which the lightest neutralino $(\tilde{\chi}_1^0)$ decays almost exclusively into a photon (γ) and a weakly interacting, stable gravitino (\tilde{G}) that gives rise to E_T by leaving the detector without depositing any energy [6]. In 13 these models, the $\tilde{\chi}_1^0$ is favored to have a lifetime on the order of a nanosecond, and the \tilde{G} is a warm dark matter 14 candidate with a mass in the range 0.5 keV $< m_{\tilde{G}} < 1.5$ keV [7]. Other direct searches [8–10] have constrained the mass of the $\tilde{\chi}_1^0$ to $\gtrsim 100~{\rm GeV/c^2}$ for much of the parameter space. At the Tevatron, sparticle production is dominated by gaugino pairs, and the $\tilde{\chi}_1^0$ mass $(m_{\tilde{\chi}_1^0})$ and lifetime $(\tau_{\tilde{\chi}_1^0})$ are the two most important parameters in determining the final states and their kinematics. Different search strategies are required for $\tilde{\chi}_1^0$ lifetimes above and below about a 18 nanosecond [11]. 19 In this letter we describe a search for GMSB with $\tau_{\widetilde{\chi}_1^0} \leq 2$ ns, favored for large $m_{\widetilde{\chi}_1^0}$, in which gaugino pairs are 20 produced and decay to the $\gamma\gamma + E_T + X$ final state where X denote other high- E_T final state particles. We use a dataset corresponding to an integrated luminosity of $2.6\pm0.2~{\rm fb^{-1}}$ of $p\bar{p}$ collisions at $\sqrt{s}=1.96~{\rm TeV}$ from the Tevatron 22 collected with the CDF II detector [12]. This is ten times more data than our previous search [8]. For the first time in 23

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Our strategy is to select $\gamma\gamma$ candidates and optimize the search for the presence of both significant E_T and large total event transverse energy to indicate the decays of heavy gauginos. We perform an *a priori* analysis based on the expected sensitivity, taking into account the background and signal predictions.

Here we briefly describe the aspects of the detector [12] relevant to this analysis. The magnetic spectrometer consists of tracking devices inside a 3-m diameter, 5-m long superconducting solenoid magnet that operates at 1.4 T.

consists of tracking devices inside a 3-m diameter, 5-m long superconducting solenoid magnet that operates at 1.4 T. A 3.1-m long drift chamber (COT) with 96 layers of sense wires measures the z position and time of the $p\bar{p}$ interaction and the momenta of charged particles. The calorimeter consists of projective towers with electromagnetic (EM) and hadronic (HAD) compartments and is divided into a central barrel that surrounds the solenoid coil ($|\eta|<1.1$) [2] and a pair of end-plugs that cover the region $1.1<|\eta|<3.6$. Both are used to identify and measure the 4-momenta of photons, electrons, jets (j) [15] and provide E_T information. The EM calorimeter is instrumented with a timing system, EMTiming [13], that measures the arrival time of photons.

Our analysis begins with events passing the CDF three-level trigger. The combined trigger selection efficiency is 12 effectively 100% for our diphoton events [14]. We then select events with two photons with $|\eta| < 1.1$ and $E_T > 13$ GeV. 13 Offline, both photons are required to be in the fiducial part of the calorimeter and to pass the standard CDF photon identification (ID) and isolation requirements [8] with two minor modifications to remove instrumental and electron 15 backgrounds [14, 16]. These events are dominated by SM production of $\gamma\gamma$, γj with $j\rightarrow\gamma_{fake}$ and $jj\rightarrow\gamma_{fake}\gamma_{fake}$, where γ_{fake} is a jet misidentified as a photon. To minimize the number of these events with large E_T due to calorimeter energy mismeasurements we remove events if the $\not\!E_{\rm T}$ vector points within $|\Delta\phi|<0.3$ of the second highest $E_{\rm T}$ photon 18 or any jet pointing to the calorimeter cracks [14]. To help maintain the projective nature of the photon reconstruction in the calorimeter we require a vertex with $|z_{\text{vertex}}| < 60 \text{ cm}$, which also helps to reduce non-collision backgrounds. For 20 events with multiple reconstructed vertices we pick the vertex with the highest $\sum_{\text{tracks}} p_T$ [2], unless assigning the photons to a different vertex lowers the E_T , in which case we choose that vertex. 22

Non-collision backgrounds coming from cosmic rays and beam-related effects can produce $\gamma\gamma + E_T$ candidates, and are removed from the inclusive $\gamma\gamma$ sample using a number of techniques. Photon candidates from cosmic rays are not correlated in time with collisions and events are removed if the timing of either photon, corrected for average path length (t_{γ}) , indicates a non-collision source [14, 16]. Photon candidates can also be produced by beam-related muons that originate upstream of the detector. These are suppressed using standard beam halo identification requirements [16]. A total of 38,053 inclusive $\gamma\gamma$ candidate events pass these requirements.

Backgrounds to the $\gamma\gamma + E_T$ final state from SM $\gamma\gamma/\gamma_{fake}/\gamma_{fake}\gamma_{fake}$ and mismeasured ("fake") E_T arise due to energy mismeasurements in the calorimeter or event reconstruction pathologies. We use METMODEL [14] to select events with real and significant E_T , as part of the optimization, and to predict the contribution of SM backgrounds with fake E_T due to normal energy measurement fluctuations. This algorithm considers the clustered (jets) and unclustered energy in the event and calculates a probability for fluctuations in the energy measurement to produce

 E_T^{fluct} equivalent to or larger than the measured E_T ($P_{E_T^{fluct} \geq E_T}$). This probability is then used to define a E_T^{fluct} significance as $-\log_{10}\left(P_{E_T^{fluct} \geq E_T}\right)$. Events with true and fake E_T^{fluct} of the same value have, on average, different E_T^{fluct} significance. We use pseudo-experiments to estimate the expected E_T^{fluct} significance distribution for SM events with fake E_T^{fluct} , and the number of mismeasured events above a given E_T^{fluct} significance requirement. The jets and unclustered energy are smeared according to their resolution functions in the event. The systematic uncertainty in the METMODEL is dominated by the uncertainty in the resolution functions [14].

The METMODEL does not account for reconstruction pathologies in SM events with no intrinsic E_T , such as a wrong choice of the primary interaction vertex or tri-photon events with a lost photon. To obtain the prediction for this background we model SM kinematics and event reconstruction using a PYTHIA Monte Carlo (MC) [17] $\gamma\gamma$ sample with a detector simulation [18]. This sample is normalized to the number of events in the inclusive $\gamma\gamma$ data sample to take into account γj and jj contributions to the backgrounds. We subtract the expectations for energy mismeasurement fluctuations in the MC to avoid double counting. Uncertainties are dominated by the statistics of the MC sample.

Electroweak production of W's and Z's which decay to leptons can also give rise to the $\gamma\gamma+E_T$ signature where one or more of the photons can be fake, but the E_T is due to one or more neutrinos. To estimate the contribution from these backgrounds we use MC simulations, normalized to their theoretical cross sections and taking into account all the leptonic decay modes of the bosons. The Baur MC [19] is used to simulate $W\gamma$ and $Z\gamma$ production and decay where initial and final state radiation (ISR/FSR) simulates $W/Z+\gamma\gamma$ events. The PYTHIA [17] MC is used to simulate backgrounds where both photons are fakes: W, Z, with photons from ISR/FSR removed, and $t\bar{t}$ sources. To minimize the dependence of our predictions on potential MC-Data differences we scale our MC predictions to the observed number of $e\gamma$ events [14] in data where we use the same diphoton triggers and analysis selection procedures developed for our inclusive $\gamma\gamma$ sample. Uncertainties are dominated by the statistics of the MC and $e\gamma$ normalization data sample.

Non-collision backgrounds are estimated using the data. Using the sample selection requirements, but requiring one of the photons to have $t_{\gamma}>25$ ns we identify a cosmic enhanced sample. Similarly, we select a beam related background enhanced sample. We estimate the number of these events in the signal region using the ratio of events outside the timing requirements to events inside the signal region and the measured efficiencies of the non-collision rejection requirements [14]. The uncertainties on both non-collision background estimates are dominated by the statistical uncertainty on the number of identified events. The top of Figure 1 shows the E_T -significance distribution for the inclusive $\gamma \gamma$ sample, along with the predictions for all the backgrounds.

We estimate the sensitivity to heavy, neutral particles that decay to photons using the GMSB reference model [6] in the mass-lifetime range, $75 \le m_{\widetilde{\chi}_1^0} \le 150$ GeV and $\tau_{\widetilde{\chi}_1^0} \le 2$ ns. Events from all SUSY processes [16] are simulated with PYTHIA [17] followed by a detector simulation [18]. The fraction of $\widetilde{\chi}_1^0$ decays that occur in the detector volume, and thus the acceptance, are dependent on both the lifetime and the masses of the sparticles [11]. The total systematic uncertainty on the acceptance, after all kinematic requirements, is estimated to be 7%, dominated by the uncertainty

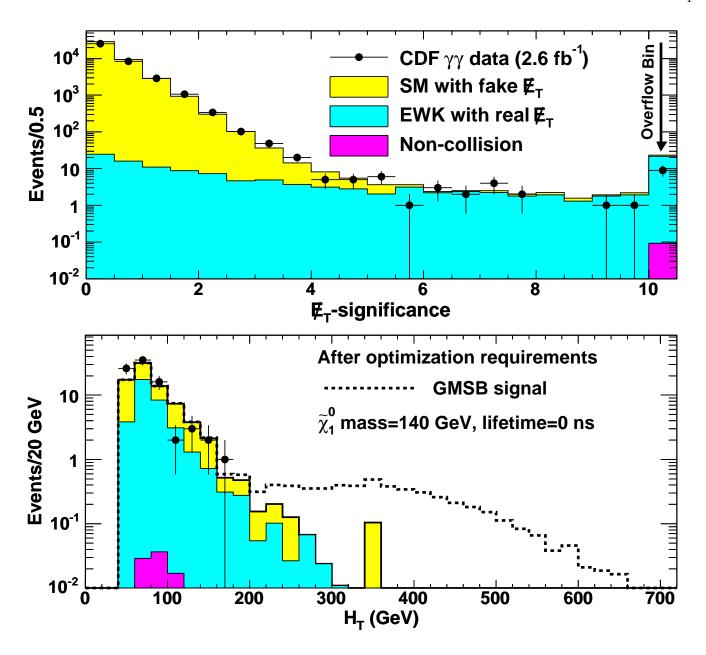


FIG. 1: The top plot shows the $\not\!E_T$ -significance distribution for the inclusive $\gamma\gamma$ candidate sample, along with the background predictions. The bottom plot shows the predicted H_T (total E_T of photons, jets and $\not\!E_T$) distribution after all but the final H_T requirement.

- in the photon ID efficiency (2.5% per photon). Other significant contributions come from uncertainties on ISR/FSR
- 2 (4%), jet energy measurement (2%), E_T-significance parameterizations (1%) and the parton distribution functions
- з (1%).
- We determine the final kinematic selection requirements by optimizing the mean expected 95% confidence level
- (C.L.) cross section limit with the no-signal assumption without looking at the data in the signal region [20]. To
- 6 compute the predicted cross section upper limit we combine the luminosity, the acceptance and the background
- ⁷ estimates with their systematic uncertainties using a Bayesian method [21]. The predicted limits are optimized by

simultaneously varying the selection requirements for E_T -significance, H_T (scalar sum of E_T of photons, jets and E_T), and the azimuthal angle between the two leading photons, $\Delta\phi(\gamma_1, \gamma_2)$. The large E_T -significance requirement eliminates most of the SM background with fake E_T . GMSB production is dominated by heavy gaugino pairs which decay to high E_T , light final state particles via cascade decays. GMSB signal has, on average, larger H_T compared to SM backgrounds so that an H_T requirement can remove these backgrounds effectively. Electroweak backgrounds with large H_T typically consist of a high E_T photon recoiling against $W \to e\nu$, identified as $\gamma_{fake} E_T$, which means the gauge boson decay is highly boosted. Thus, the two photon candidates in the final state are mostly back-to-back. Also, the high E_T diphotons with large H_T from SM background are mostly back-to-back with fake E_T ; the $\Delta\phi(\gamma_1, \gamma_2)$ cut, therefore, reduces both these backgrounds.

While each point in the considered $\tau_{\Sigma 0}$ vs. $m_{\Sigma 0}$ space gives a slightly different optimization, we choose a single

While each point in the considered $\tau_{\widetilde{\chi}_1^0}$ vs. $m_{\widetilde{\chi}_1^0}$ space gives a slightly different optimization, we choose a single set of requirements to maximize the expected 95% C.L. exclusion region; where the predicted production cross section at next-to-leading order [22] is above the expected cross section limit. The exclusion region also takes into account the production cross section uncertainties which are dominated by the parton distribution functions (7%) and the renormalization scale (3%). We find the optimal set of cuts, before unblinding the signal region, to be: $E_{\rm T} = \frac{E_{\rm T}}{E_{\rm T}} = \frac{E_{\rm T}}{E_{\rm$

No events in the data pass the final event selection. The predicted $H_{\rm T}$ distribution is shown in Figure 1 (bottom), after all but the final $H_{\rm T}$ cut. The data is consistent with the no-signal hypothesis and well modeled by backgrounds alone. We set cross section limits as a function of $m_{\widetilde{\chi}_1^0}$ and $\tau_{\widetilde{\chi}_1^0}$ respectively, as shown in Figure 2. The $m_{\widetilde{\chi}_1^0}$ reach, based on the predicted and observed number of events for $\tau_{\widetilde{\chi}_1^0} = 0$, is 141 GeV/c² and 149 GeV/c² respectively. These limits significantly extend the search sensitivity beyond the results of D0 [9] for $\tau_{\widetilde{\chi}_1^0} \le 2$ ns and, when combined with the complementary exclusions from CDF and LEP [10, 16], cover the region shown in Figure 3.

In conclusion, we have performed an optimized search for heavy, neutral particles that decay to photons in the $\gamma\gamma+E_T$ final state using 2.6 fb⁻¹ of data. There is no excess of events beyond expectations. We set cross section limits using a GMSB model with $\tilde{\chi}_1^0 \to \gamma \tilde{G}$, and find an exclusion region in the $\tau_{\tilde{\chi}_1^0}$ - $m_{\tilde{\chi}_1^0}$ plane with the world-best 95% C.L. lower limit on the $\tilde{\chi}_1^0$ mass of 149 GeV/c² at $\tau_{\tilde{\chi}_1^0}$ =0 ns. By the end of Run II, an integrated luminosity of 10 fb⁻¹ is possible for which we estimate a mass reach of $\simeq 160 \text{ GeV/c}^2$ at a lifetime of 0 ns.

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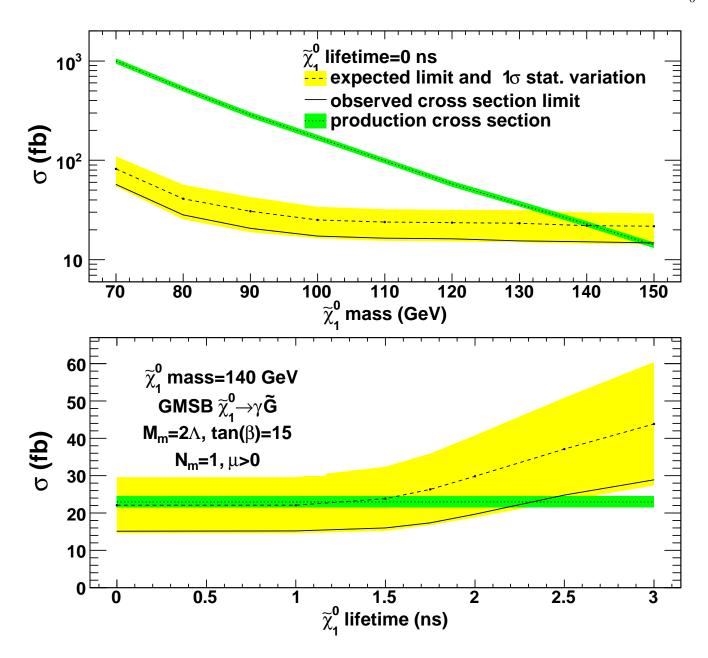


FIG. 2: The predicted and observed 95% C.L. cross section upper limits as a function of the $\tilde{\chi}^0_1$ mass at a lifetime of 0 ns (top) and as a function of the $\tilde{\chi}^0_1$ lifetime at a mass of 140 GeV/ c^2 (bottom). Indicated in green (darker shading) is the production cross section, along with its 8.0% uncertainty-band. In yellow (lighter shading) is the RMS variation on the expected cross section limit.

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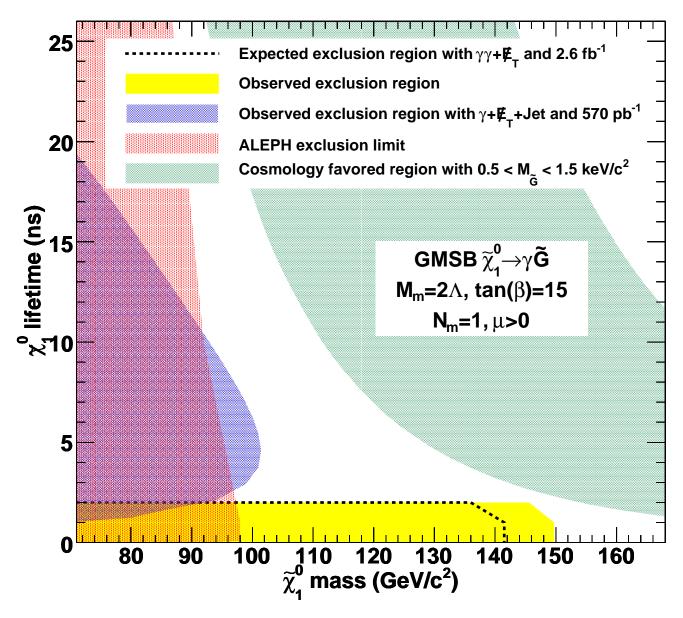


FIG. 3: The predicted and observed exclusion region along with the limit from ALEPH/LEP [10] and the CDF $\gamma + E_T + jet$ 'delayed' photon analysis [16]. We have a mass reach of 141 GeV/ c^2 (predicted) and 149 GeV/ c^2 (observed) at lifetimes up to 1 ns. The shaded band shows the parameter space where $0.5 < m_{\tilde{G}} < 1.5 \text{ keV}/c^2$, favored by cosmological models [7].

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